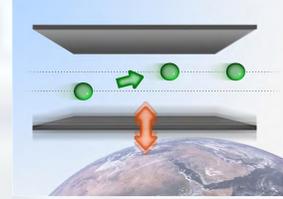


neutrons: what for? and how?

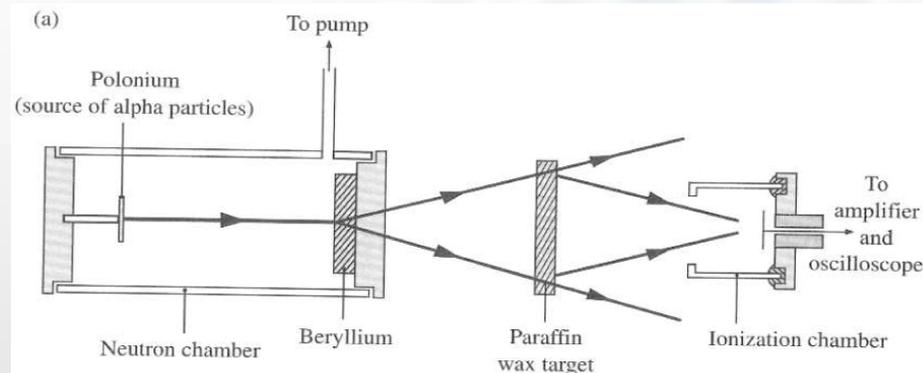
Christian Vettier
ESRF Grenoble, France
ESS AB, Lund Sweden

- neutron properties
- neutron scattering applications
- neutron sources
- instruments and applications



the discovery of the neutron

James Chadwick 1932



Cavendish Laboratory,
Cambridge

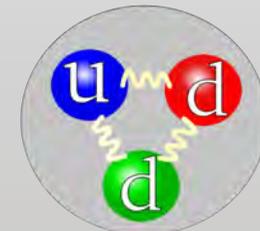
neutrons: no charge, mass close to proton – $1.675 \cdot 10^{-23}$ kg

spin 1/2 - magnetic moment $1.93 \mu_N$



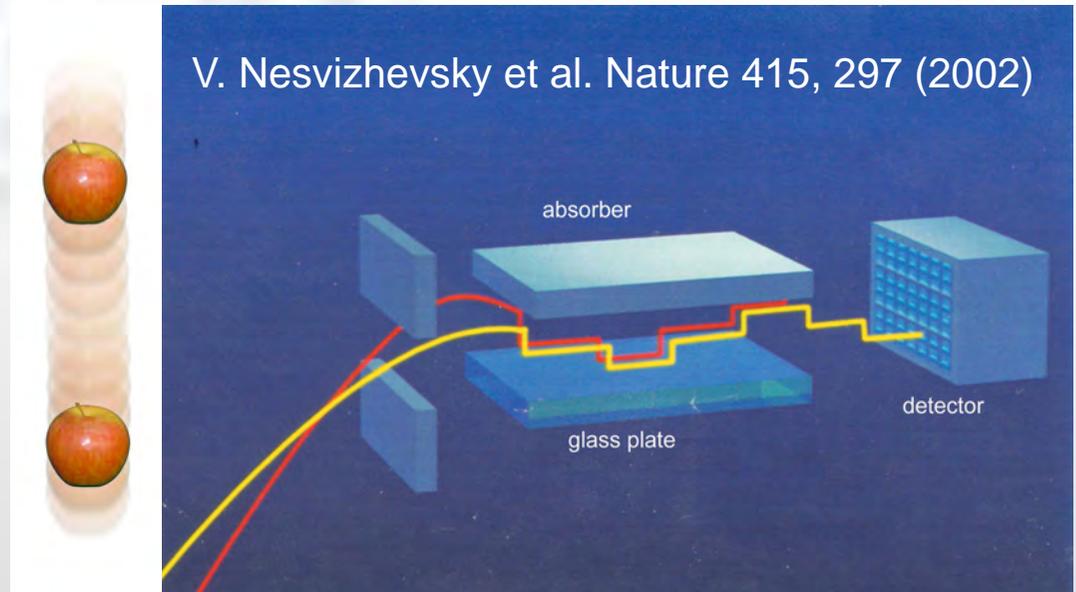
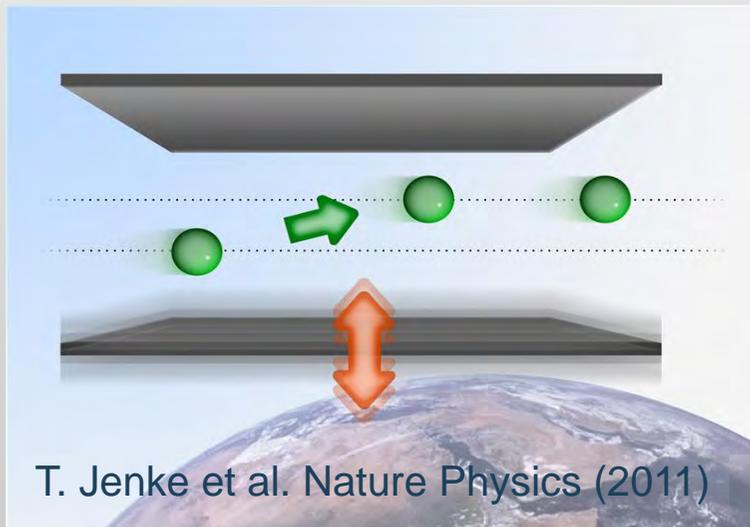
free neutrons are unstable β -decay proton, electron, anti-neutrino

life-time 886 ± 1 sec



neutron : an object to be studied
fundamental interactions

confirmation of quantum effects in
presence of gravity field



new developments
resonance spectrometry

wave-particle duality

diffraction of thermalised neutrons

kinetic energy

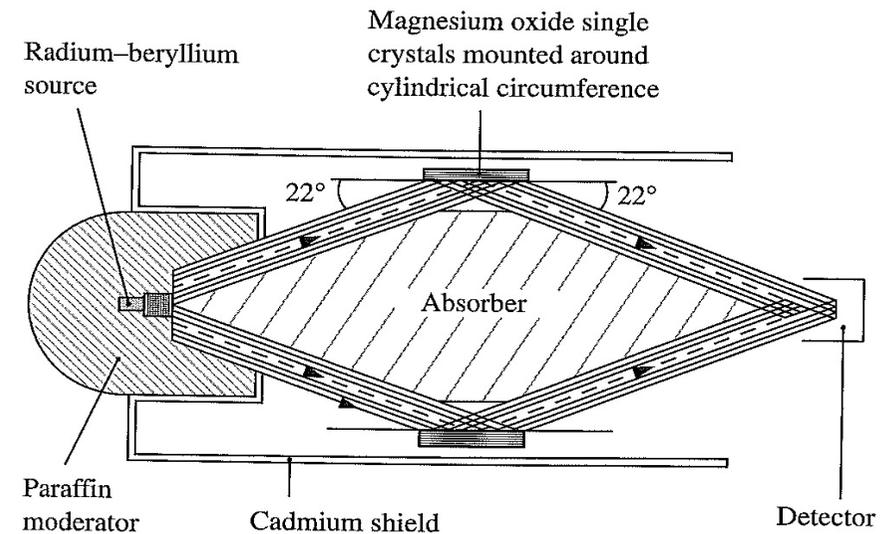
$$E(\text{meV}) = 2.072 \left[k(\text{\AA}^{-1}) \right]^2 \approx \left[\frac{9}{\lambda(\text{\AA})} \right]^2$$

$$v(\text{m/s}) = \frac{3,956}{\lambda(\text{\AA})}$$

$$T = \frac{L}{v} = 252.77 \mu\text{sec} \cdot \lambda[\text{\AA}] \cdot L[\text{m}] \quad \text{fortunately large value!}$$

monochromatisation: diffraction or velocity selection

neutron $dh = 1 \text{ m} \approx 0.1 \mu\text{eV}$ in gravity on Earth



D.P Mitchell and N. Powers (1936)

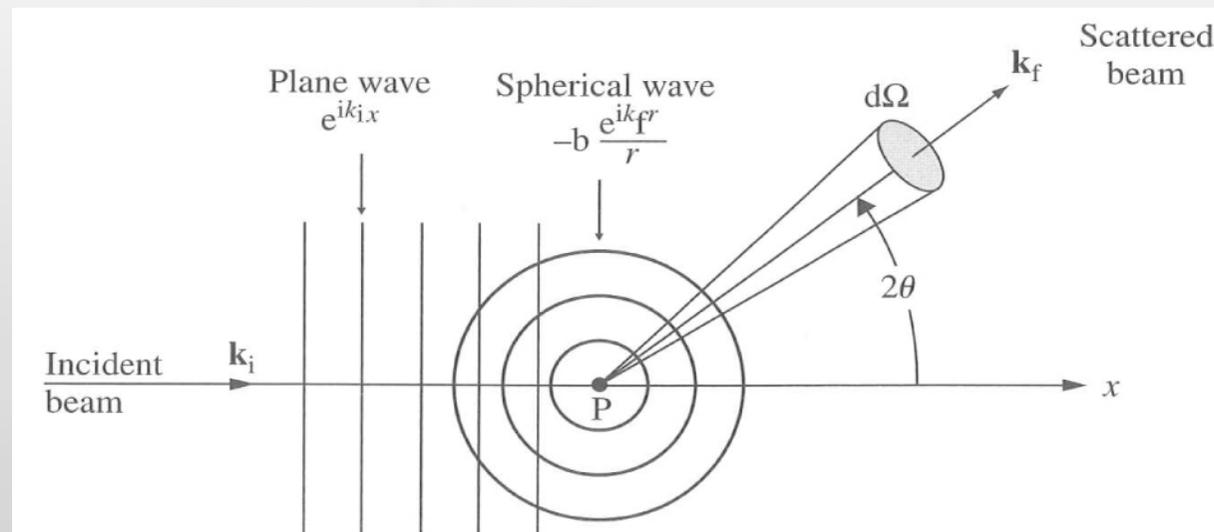
$$1(\text{\AA}) \approx 82 \text{ meV}$$

$$1 \text{ meV} \approx 8 \text{ cm}^{-1} \approx 0.25 \text{ THz} \approx 11.6 \text{ K}$$

what do neutrons ‘feel’ or ‘interact’ with?

like a billiards ball and a small compass

single nucleus process



neutron-nuclei interaction

Fermi length $b \approx 10^{-12}$ cm

magnetism

magnetisation densities

magnetic interaction



spin 1/2 total magnetic moment

$$M \approx \langle L \rangle + 2\langle S \rangle$$

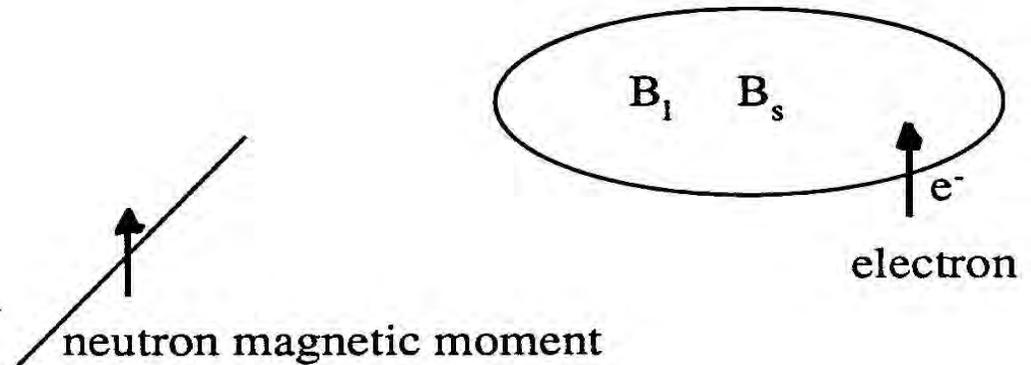
order of magnitude $g_n r_0 M / 2 \approx 0.5 \cdot 10^{-12} \text{ cm}$

similar to x-ray Thomson scattering

$$Z r_0 \approx 10^{-12} \text{ cm}$$

neutron-nuclei interactions

Fermi length $b \approx 10^{-12} \text{ cm}$



scattering form factors

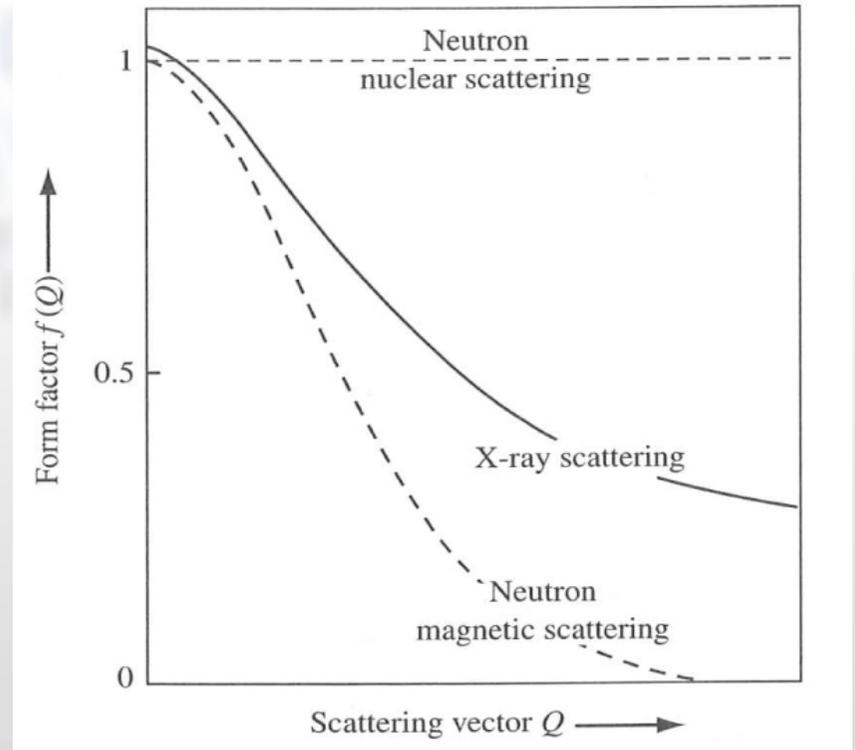
brief summary

neutron-nuclei interactions

- isotopic sensitivity
- crystal structures
- atomic motions

neutron-magnetic moments interactions

- magnetic structures
- spin dynamics magnetic interactions
- “magnetic” electron densities



coherent and incoherent scattering

$$\frac{d\sigma}{d\Omega} = \left| \sum_j b_j \exp(i\mathbf{Q} \cdot \mathbf{r}_j) \right|^2 \quad \text{different sites, no correlation} \quad \langle b_j b_{j'} \rangle = \langle b_j \rangle \langle b_{j'} \rangle = \langle b \rangle^2$$

$$\frac{d\sigma}{d\Omega} = \sum_{j,j'} \langle b \rangle^2 \exp[i\mathbf{Q} \cdot (\mathbf{r}_j - \mathbf{r}_{j'})] + \sum_j (\langle b^2 \rangle - \langle b \rangle^2) \quad \text{same site} \quad \langle b_j b_{j'} \rangle = \langle b_j^2 \rangle = \langle b^2 \rangle$$

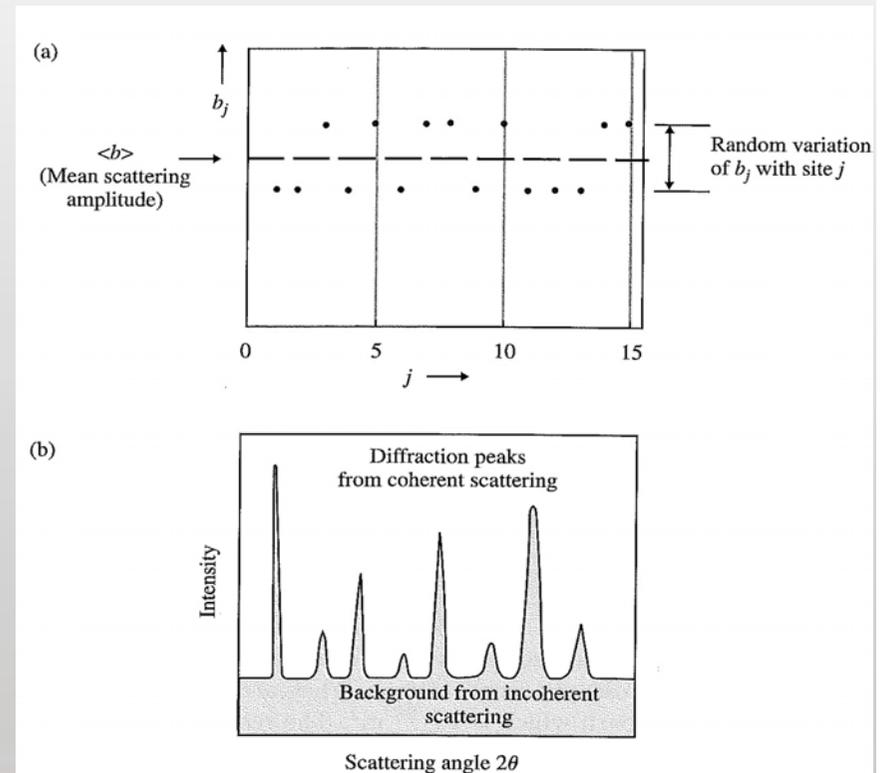
$$\sigma_{\text{coh}} = 4\pi \langle b \rangle^2 = 4\pi b_{\text{coh}}^2 \quad \sigma_{\text{incoh}} = 4\pi (\langle b^2 \rangle - \langle b \rangle^2)$$

isotopic dependence

nuclear spin dependence, I

$$\sigma_{\text{coh}} = 4\pi (\omega_+ b_+ + \omega_- b_-)^2$$

$$\sigma_{\text{incoh}} = 4\pi \omega_+ \omega_- (b_+ - b_-)^2 \quad \omega_+ = \frac{I+1}{2I+1} \quad \omega_- = \frac{I}{2I+1}$$



what do we observe/measure with neutrons ?

isotope labelling (deuteration

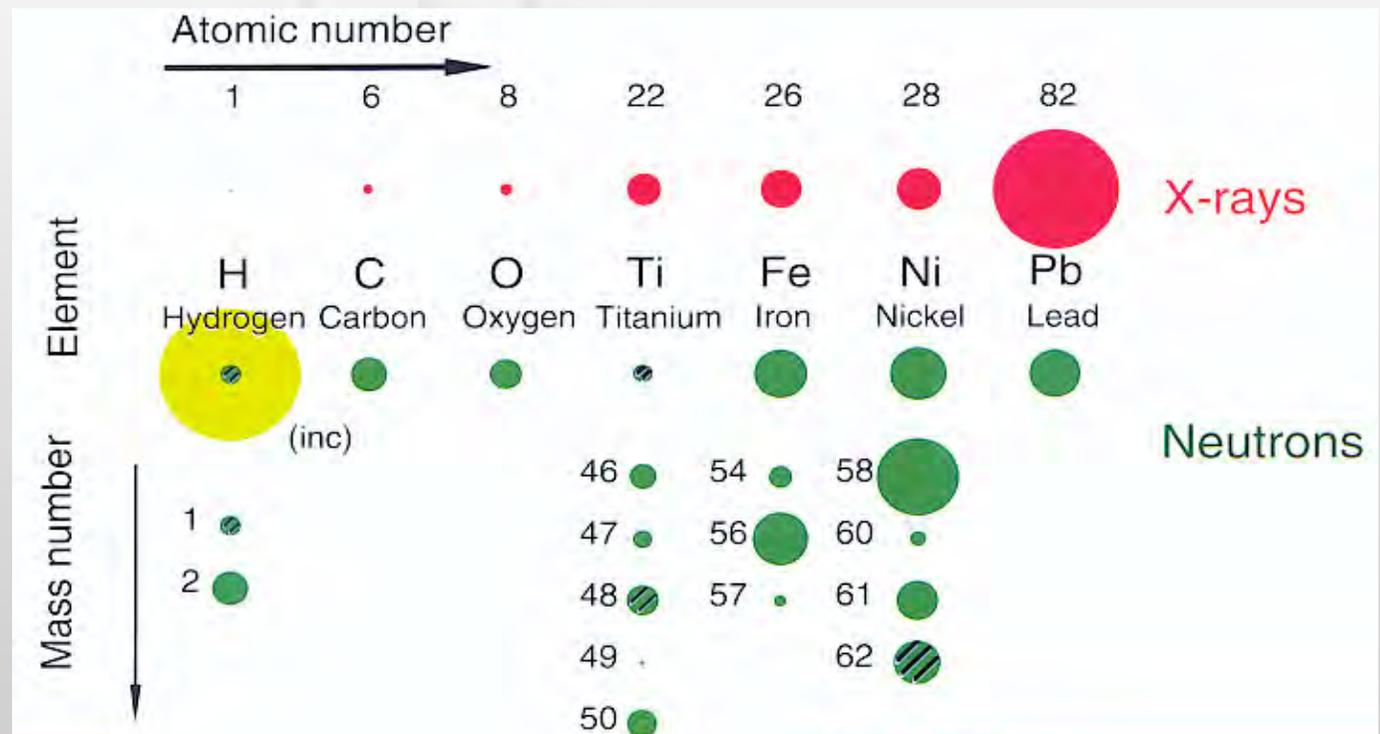
neutron-nuclei interactions
Fermi length $b \approx 10^{-12}$ cm

some b 's are < 0

... phase shift ...

isotopic sensitivity

incoherent scattering



contrast variation

objects in solution

different average cross-sections

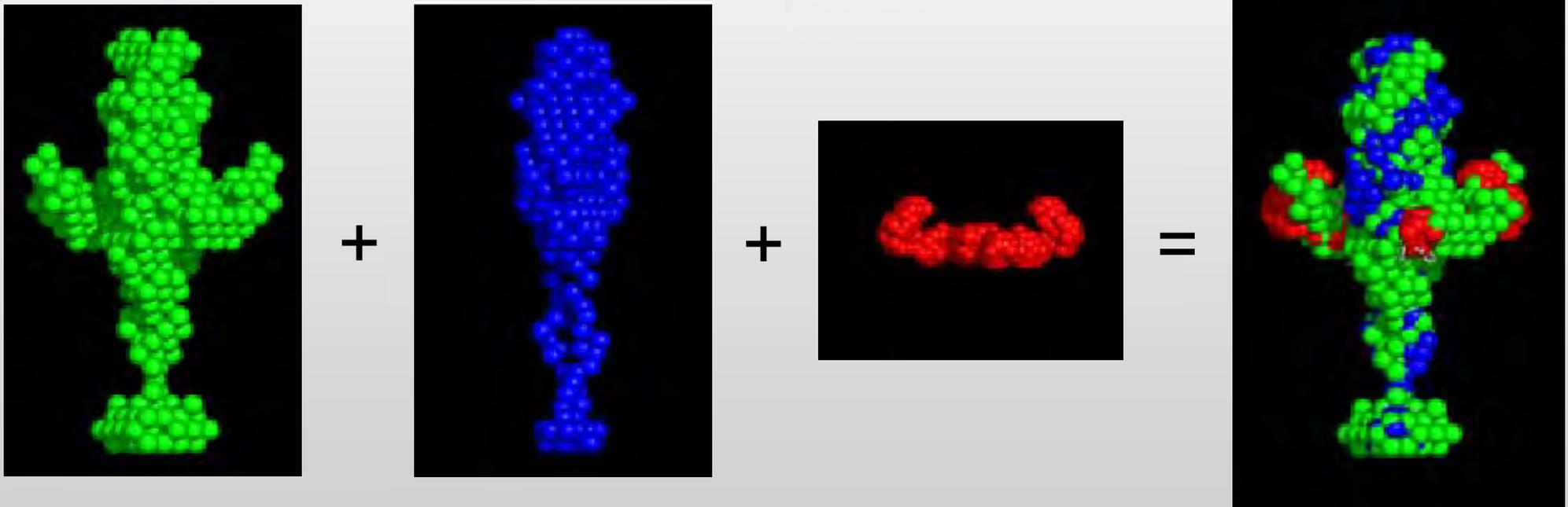


100 % H₂O

▶ 100 % D₂O

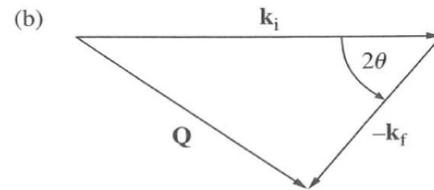
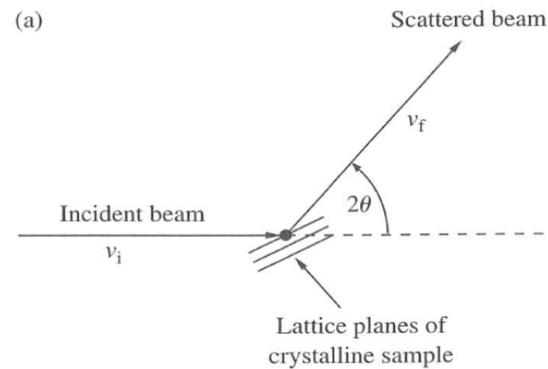
- null or full contrast : external shapes
- intermediate contrast : details

sensitivity and selectivity
isotopic substitution/contrast variation



scattering geometry

velocities



momentum space

$$\frac{\hbar^2 Q^2}{2m_n} = E_i + E_f - 2\sqrt{E_i(E_i - \hbar\omega)} \cos(2\theta)$$

kinematics

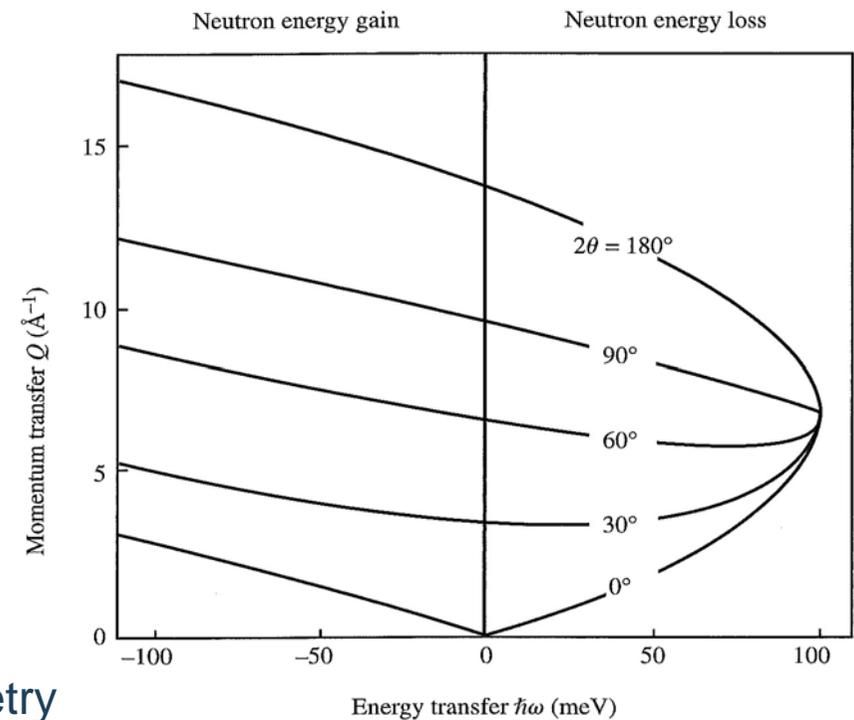
fixed E_i

direct geometry

$$Q = k_i - k_f$$

$$\hbar\omega = E_i - E_f = \frac{\hbar^2}{2m_n} (k_i^2 - k_f^2)$$

$$Q^2 = k_i^2 + k_f^2 - 2k_i k_f \cos(2\theta)$$



how to produce neutrons?

free neutrons are created through 'escape' from nuclei

neutron stars



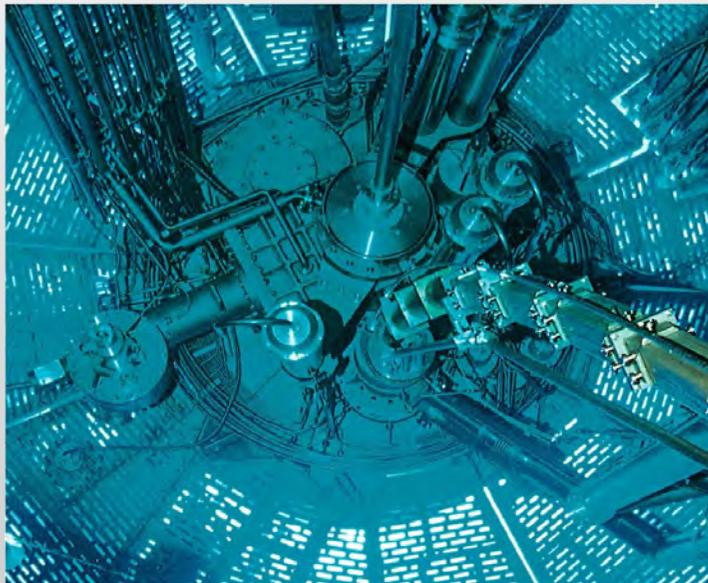
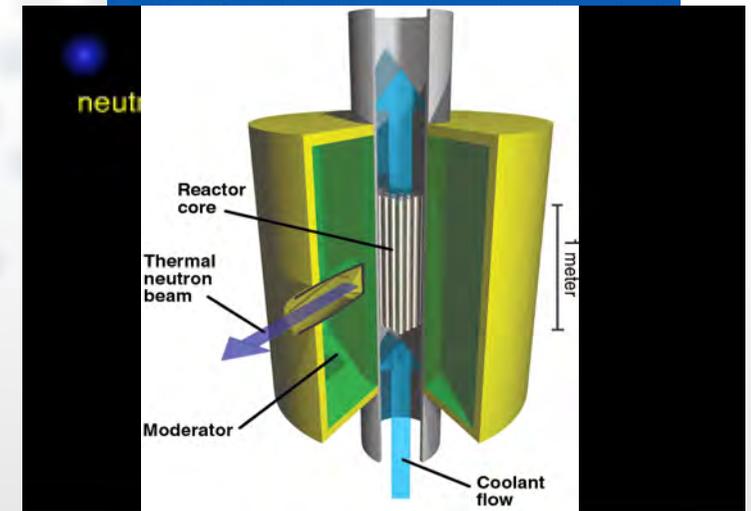
neutrons from fission process
neutrons from spallation process
neutrons from fusion process

a neutron source

how does it work?

fission reactor

slow neutrons splitting nuclei



ILL reactor

D₂O moderator

hot source: graphite block 2400K

cold sources: liquid hydrogen 25K

thermal flux $1.5 \cdot 10^{15}$ n/cm²/sec

high power density (10 MW/litre) / achievable rate of cooling reactor core

pulsed reactors

time-of-flight methods on pulsed sources

pulsing reactivity to become supercritical

- mechanically (fissile materials or reflector)
- electron pulses in the core - photo-fission process
- combination of the two above.

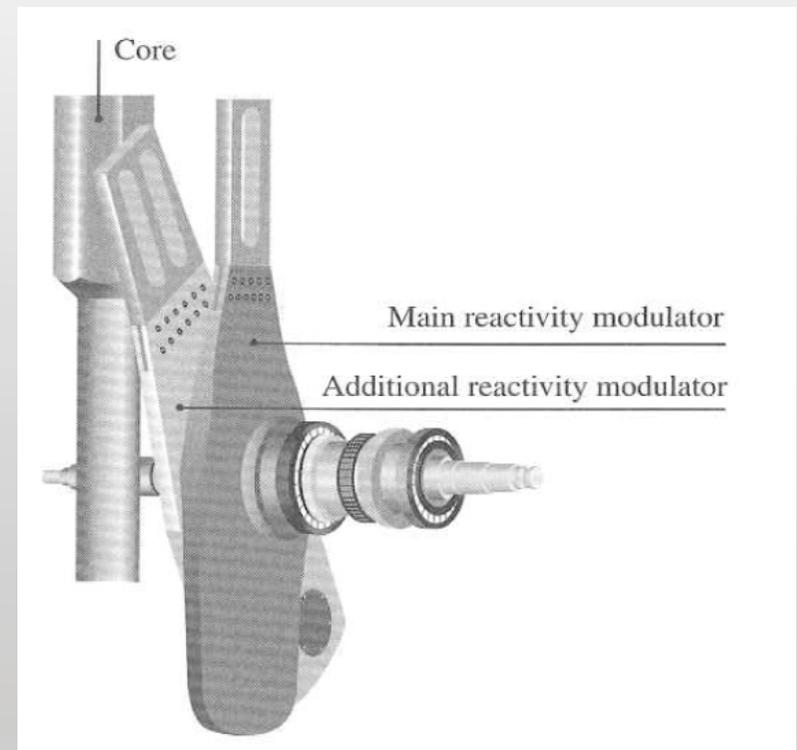
IBR-2 at Dubna, Russia

average power 2 MW

peak power 1,500 MW

rep rate 5 Hz, pulse width 215 μ sec

peak flux $1.5 \cdot 10^{16}$ n/cm²/sec



pulsed sources

to overcome cooling problems

series of installations at Argonne Nat Lab, Rutherford, KEK

photo-fission

using high-energy electrons on heavy metal targets

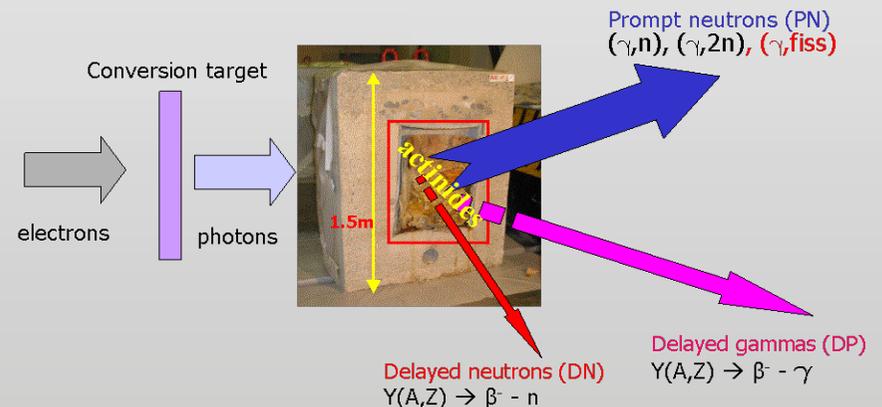
deceleration leads to Bremsstrahlung

γ -rays produce fast neutrons

not efficient

3,000MeV per neutron

intense γ -ray beams



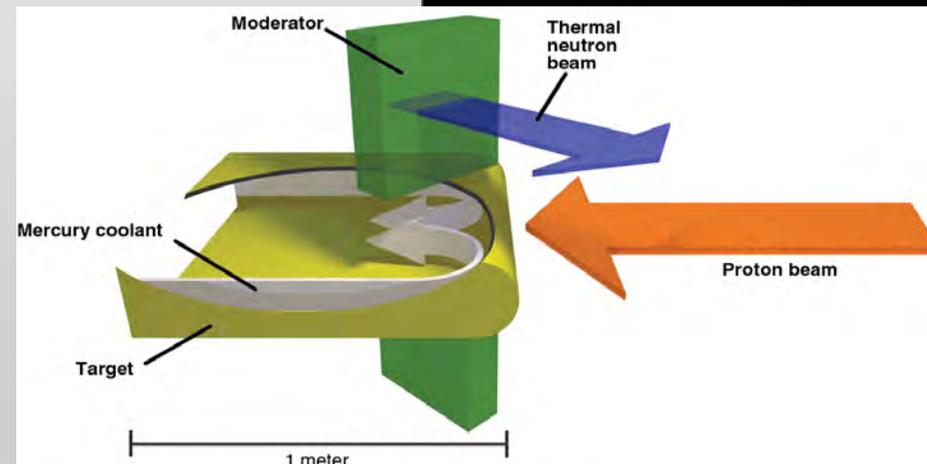
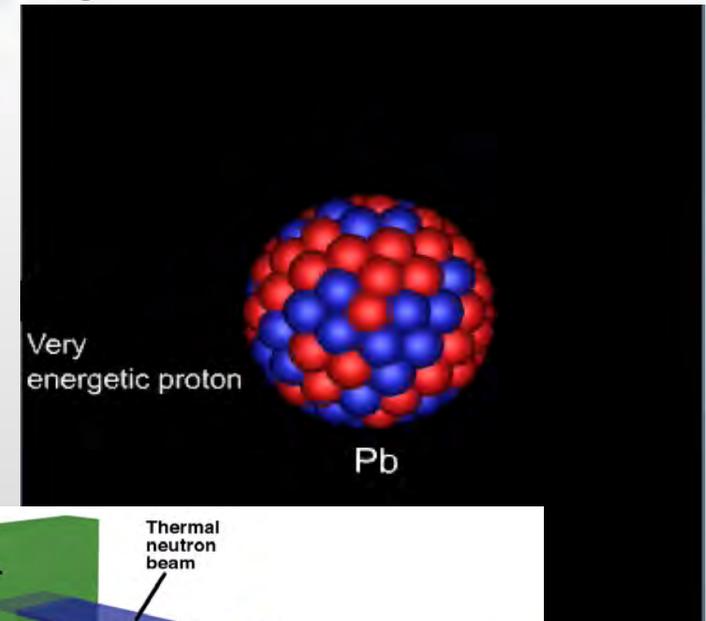
spallation sources

using high-energy electrons on heavy metal targets

‘evaporation’ of 30 fast neutrons

accelerator proton 2.0 GeV

high Z metal target



ISIS facility Rutherford Laboratory

H⁻ ion source

Linac 70 MeV

proton synchrotron 800 MeV

repetition rate 50 Hz

pulse width 100 ns

power 0.200 MW

2 target stations different moderators



neutron pulses width ~ 10-90 μ s

PSI
protons pulses 0.3ns 50.6MHz
equivalent to 0.7MW
continuous neutron source but pulsed accelerator



SNS



1 MW
short pulses

JPARC



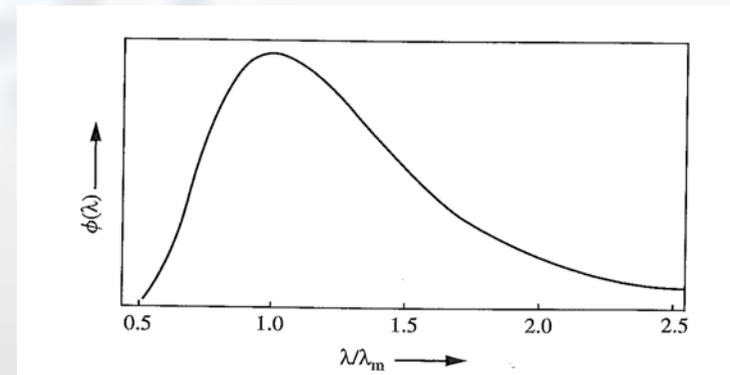
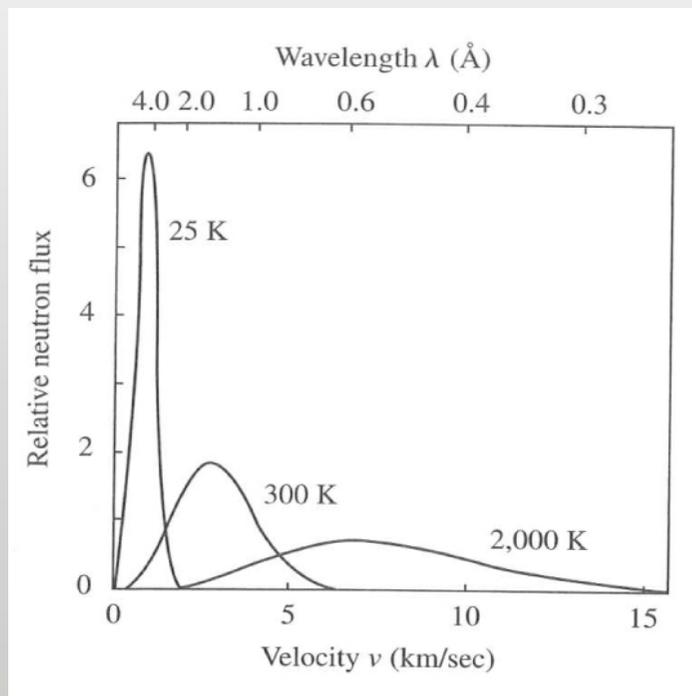
0.8 MW 60Hz
short pulses
in 2012?

fission/spallation neutrons are in the MeV range

moderation process

thermal equilibrium with moderator

Maxwellian spectrum if thermal equilibrium



$$\lambda_m = \frac{h}{(5m_n k_B T)^{1/2}} = \frac{19.483}{(T)^{1/2}} (\text{Å})$$

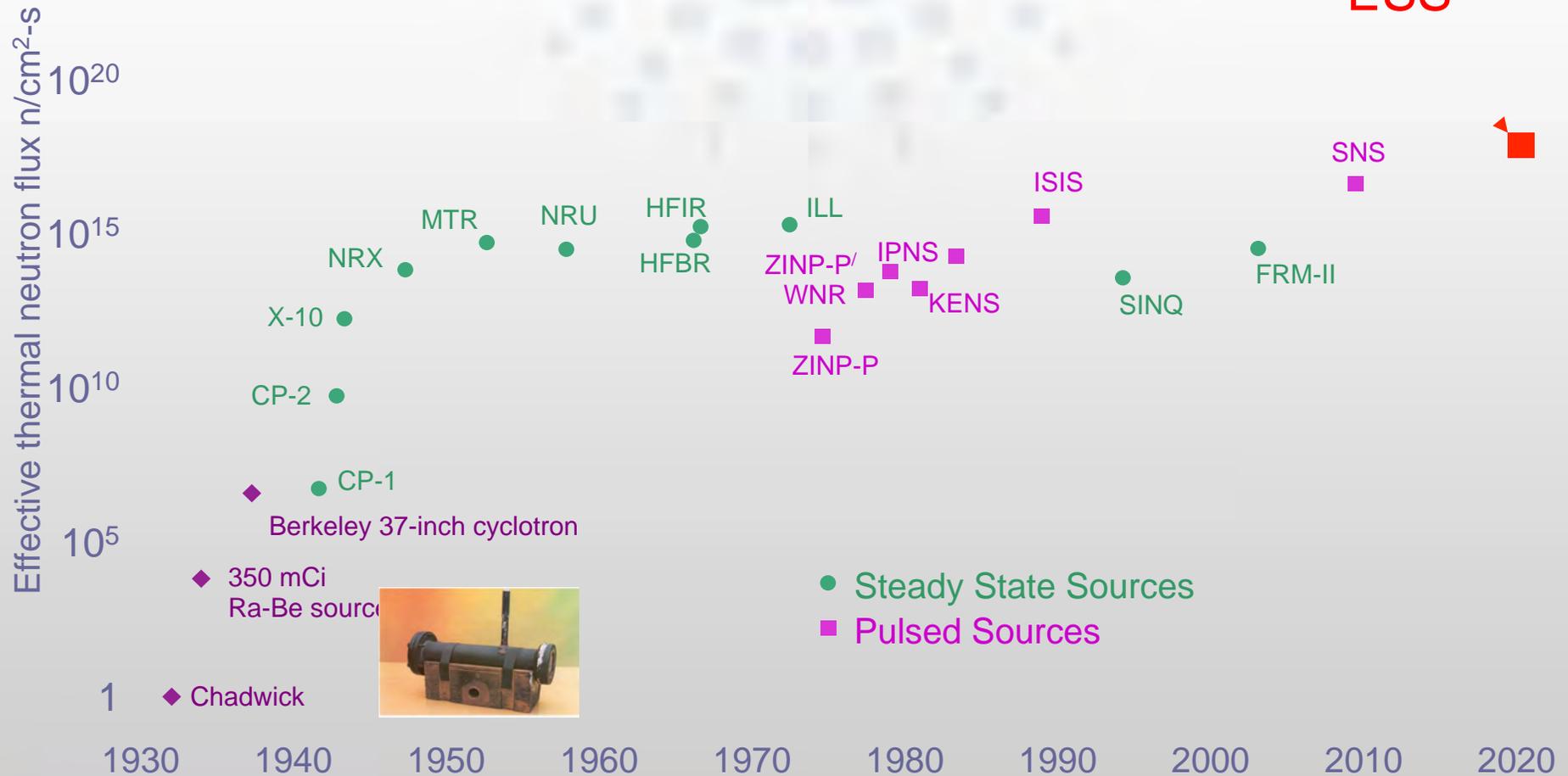
choice of moderator depends on applications

hot neutrons <0.6 Å

thermal neutrons 1- 4 Å

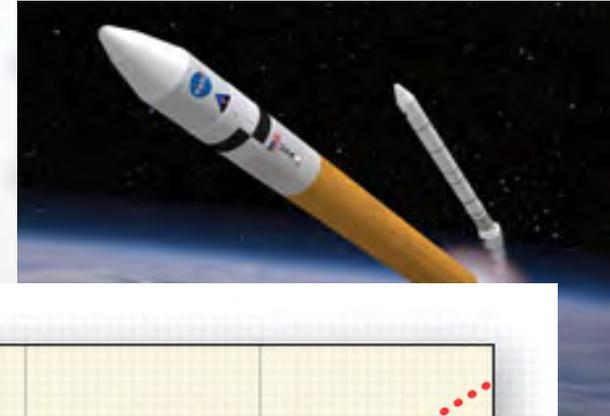
cold neutrons 4-20 Å

evolution of neutron sources characteristics

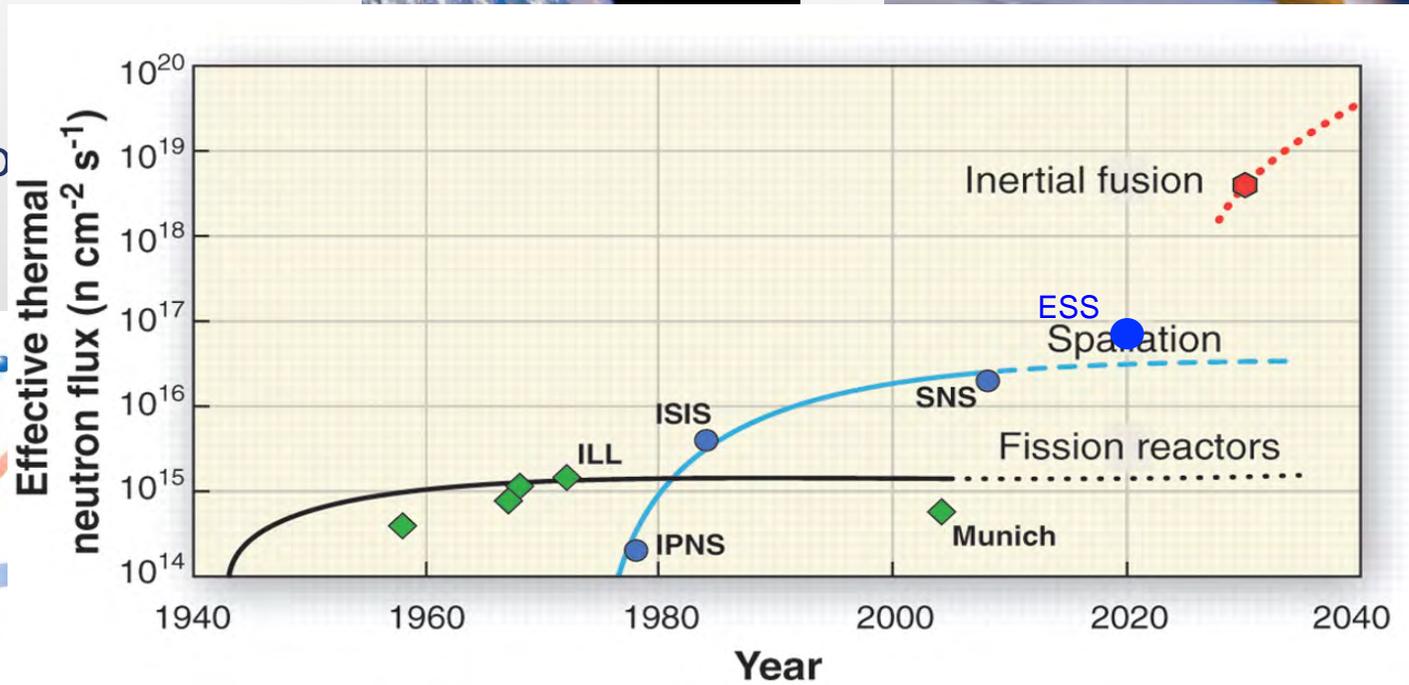
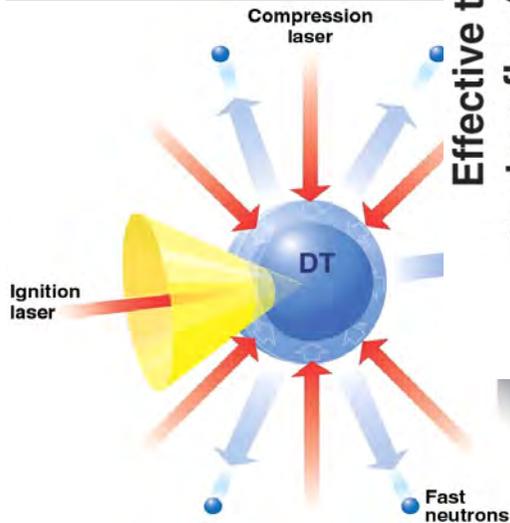


(Updated from *Neutron Scattering*, K. Skold and D. L. Price, eds., Academic Press, 1986)

landing the moon again by 2019!



a 'dream' neutron

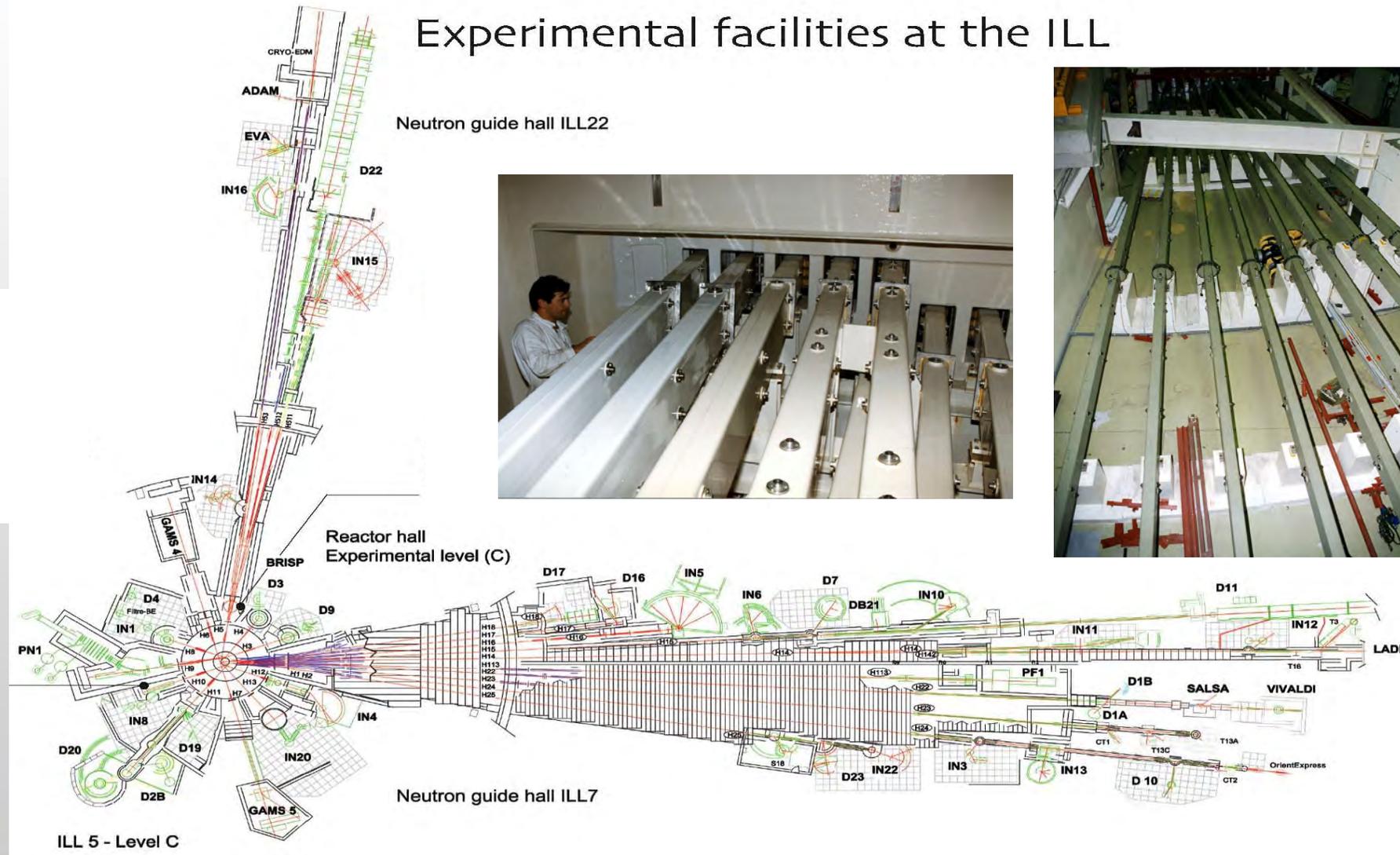


propagating thermonuclear reaction

A. Taylor et al. Science 315, 1092 (2007)

neutron guide halls

Experimental facilities at the ILL



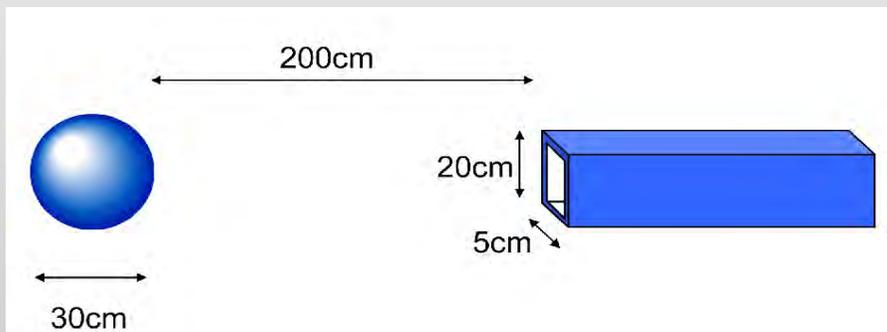
low brightness sources

$\sim 10^{14}$ n/cm²/steradian/sec

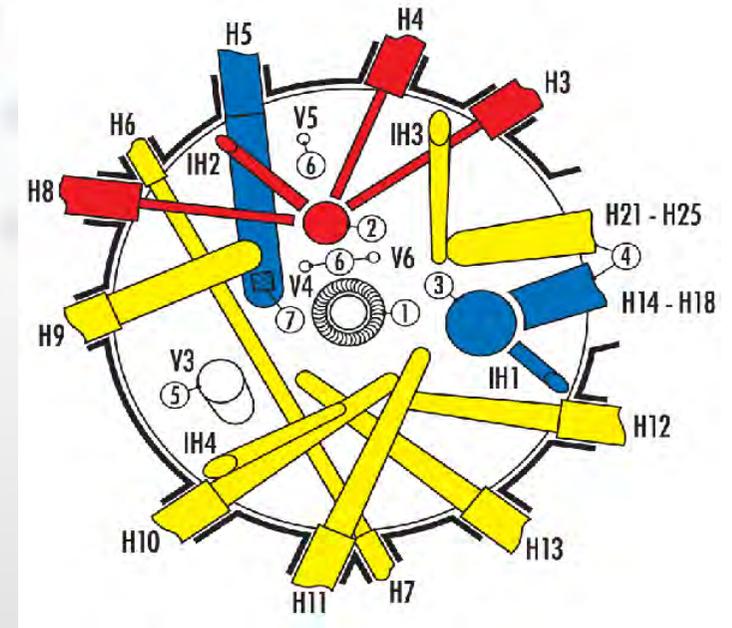
compared with a 60 W (900 lm) light bulb

$\sim 10^{19}$ photons/cm²/steradian/sec

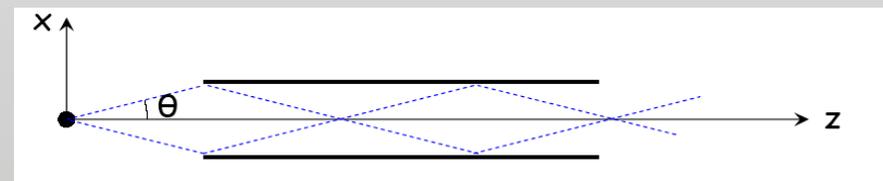
optimise neutron economy



a few degrees!



- to collect (thermal/cold) neutrons
- to transport them to instruments



crystal monochromators

selection of wavelength bandwidth (energy)

focusing

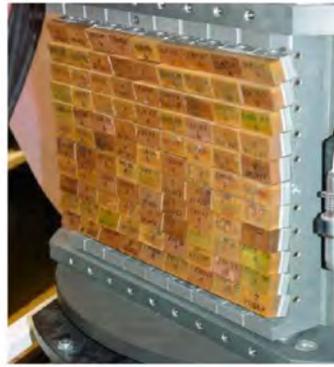
focusing mechanics



Graphite 002

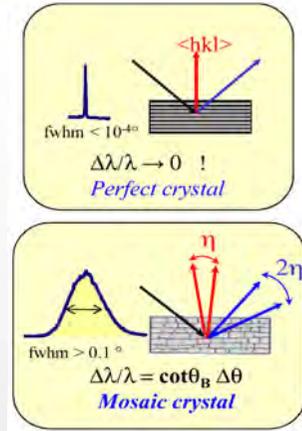


Copper 200



Heusler alloy: polarising monochromator

focusing supermirrors / guides



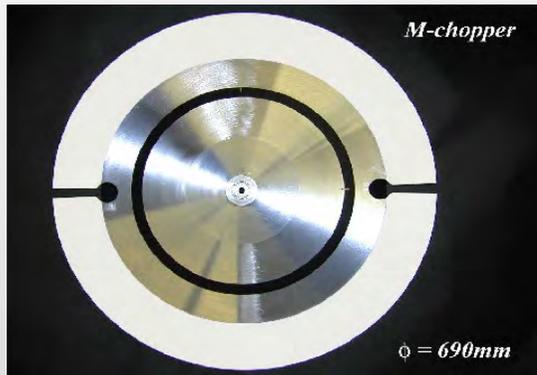
principle for time-of-flight

create pulse structure (even on reactors)

separate energies through velocity

select energies via time slicing

disk chopper

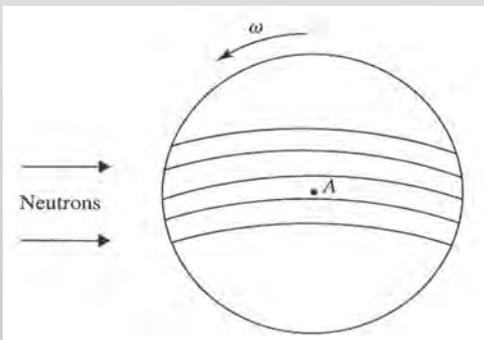


opening of ~ 1 degree

rotational speed 300Hz

opening of about 80 μsec

disk diameter of 700 cm (speed about 600m/s)



Fermi chopper

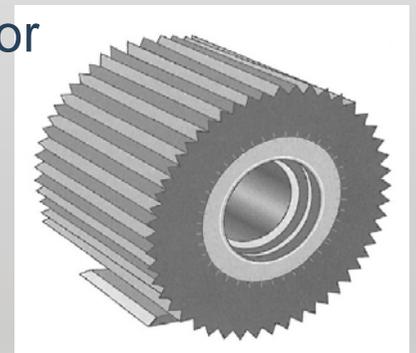
rotational speed 600Hz

opening of about 1 μsec

helical velocity selector

typical bandwidth 10%

SANS experiments

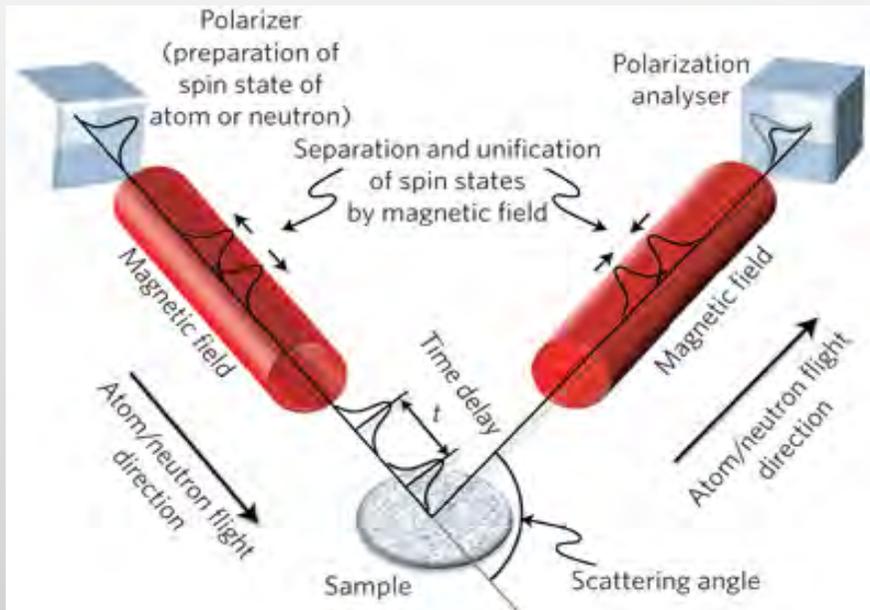


high resolution method – neutron spin echo

create neutron spin states

exploit Larmor precession of neutrons

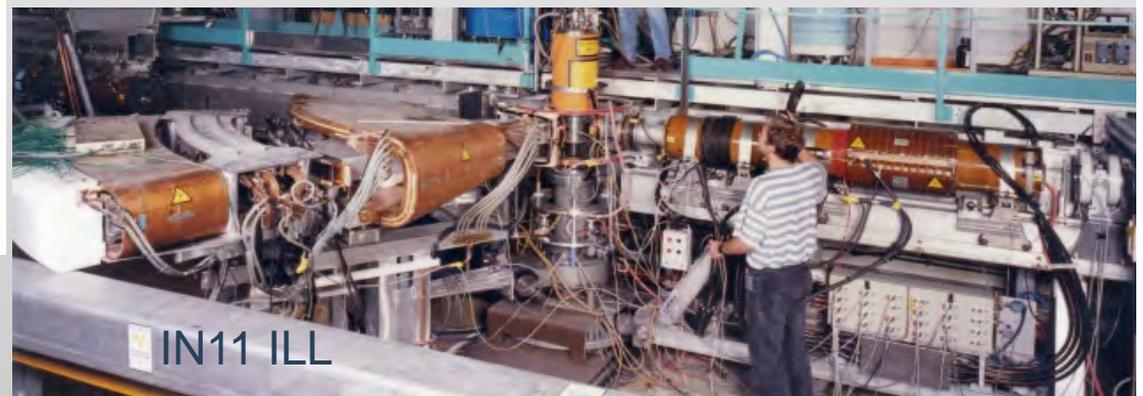
vary time shift to probe time-dependent correlations



typical time scales:

0.01 - 500 ns

400 μeV – 8 neV



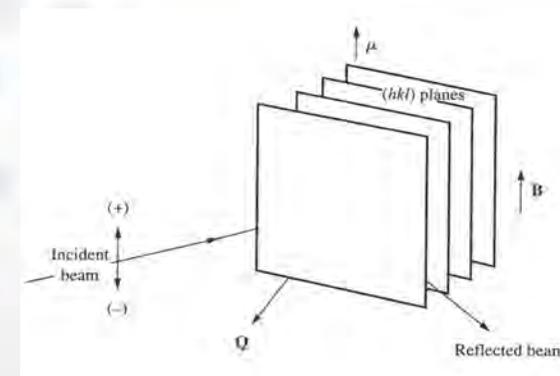
H. Hedgeland et al. Nature Phys. 5, 561 (2009)

production of polarised beams

Bragg reflection from magnetised crystal

$$\frac{d\sigma}{d\Omega} = F_N^2(\mathbf{Q}) + 2(\mathbf{P} \cdot \boldsymbol{\mu}) |F_N(\mathbf{Q})| |F_M(\mathbf{Q})| + F_M^2(\mathbf{Q})$$

B is guide field and polarising field



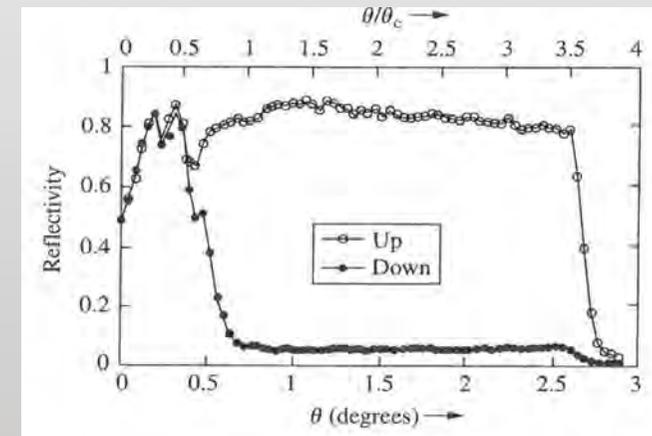
up-state $\frac{d\sigma}{d\Omega} = |F_N(\mathbf{Q}) + F_M(\mathbf{Q})|^2$ down-state $\frac{d\sigma}{d\Omega} = |F_N(\mathbf{Q}) - F_M(\mathbf{Q})|^2$
 good matching for Cu_2MnAl (111), $\text{Co}_{0.92}\text{Fe}_{0.08}$ (200), ^{57}Fe (110)

total reflection from magnetised mirror



$$n' = 1 - \frac{N\lambda^2(\bar{b} \pm p)}{2\pi}$$

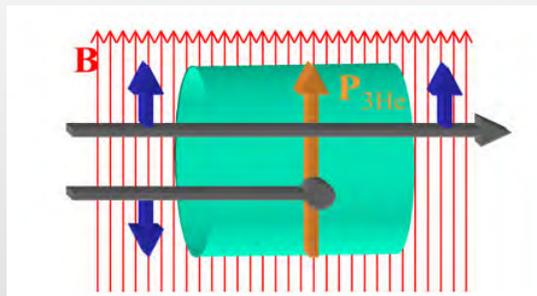
$$\theta_c \approx \lambda \left[\frac{N}{\pi} (\bar{b} \pm p) \right]^{1/2}$$



production of polarised beams

transmission through polarising filter

neutron absorption cross section of ^3He is highly spin-dependent

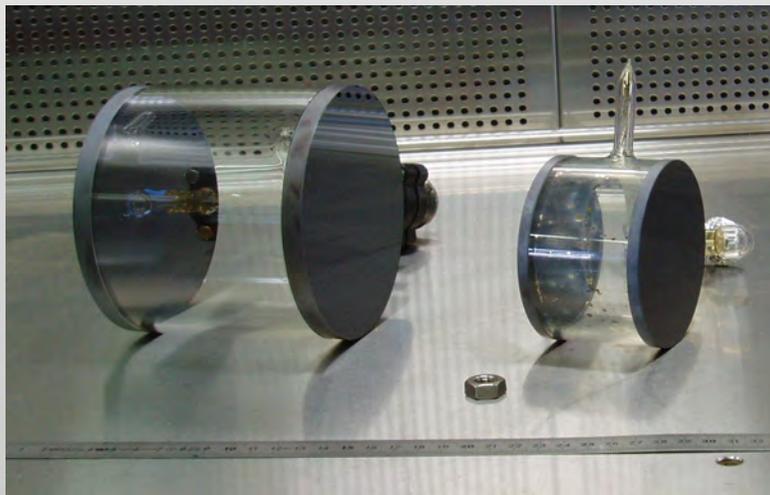


$$\sigma_{\uparrow\uparrow} \approx 0$$

parallel neutron spins are transmitted

$$\sigma_{\uparrow\downarrow} \approx 6000 \text{ barns}$$

anti-parallel neutron spins are absorbed



nuclear spin-polarised
 ^3He gas

the neutron is a weak probe! difficult to detect
 thermal neutrons produce negligible ionisation

secondary ionisation due to neutron capture

emission of gamma rays or charged nuclei

most common absorbers: ^3He , ^{10}B and ^6Li

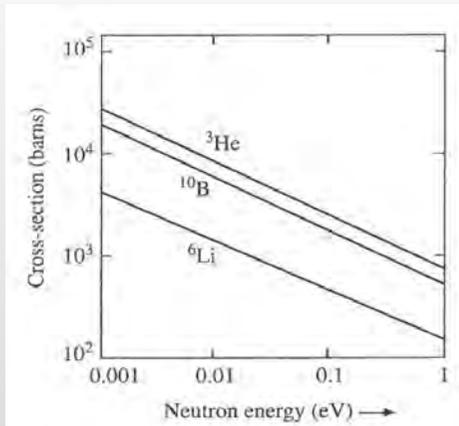
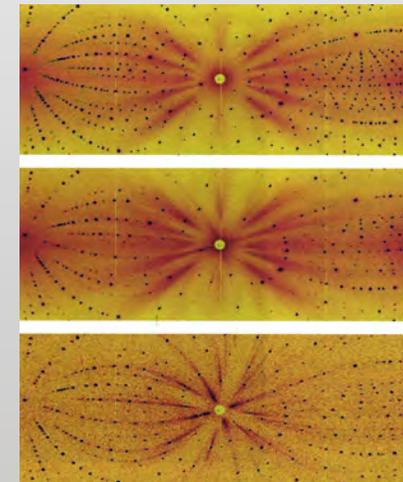


image plates

common in SRX community

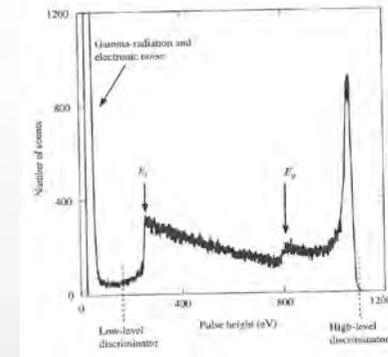
need to convert neutron into photons (^6Li or Gd)



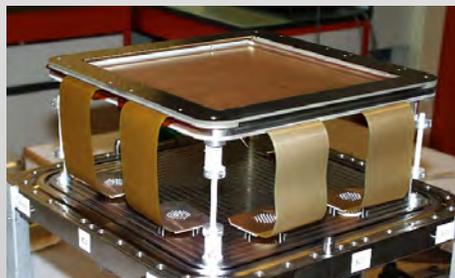
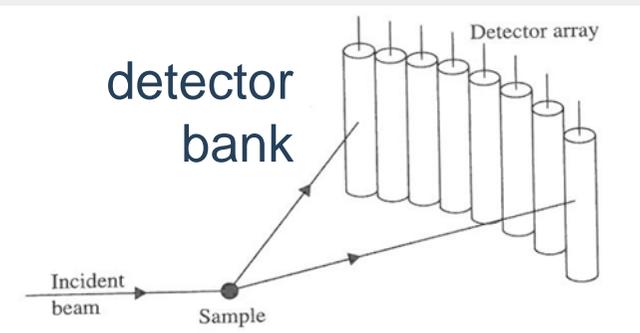
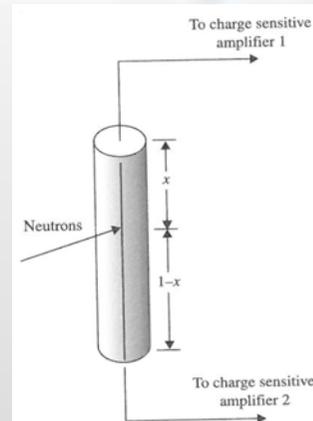
gas detectors

^3He gas difficult to obtain nowadays

^{10}B in BF_3 toxic gas



position sensitive detectors

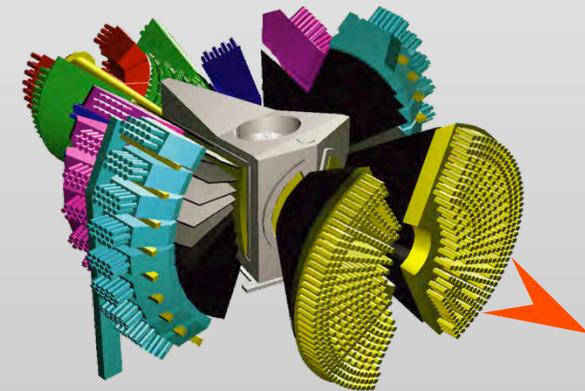
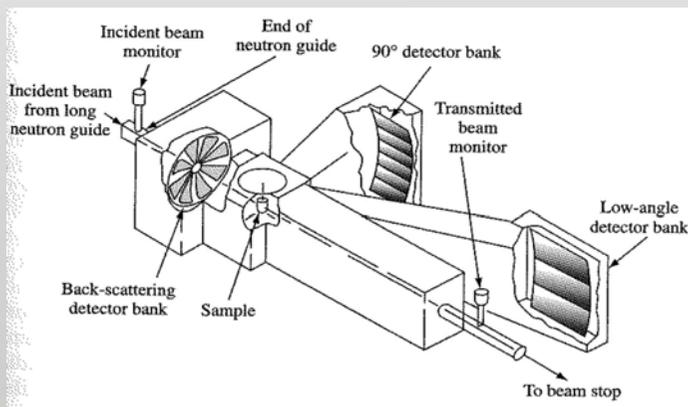
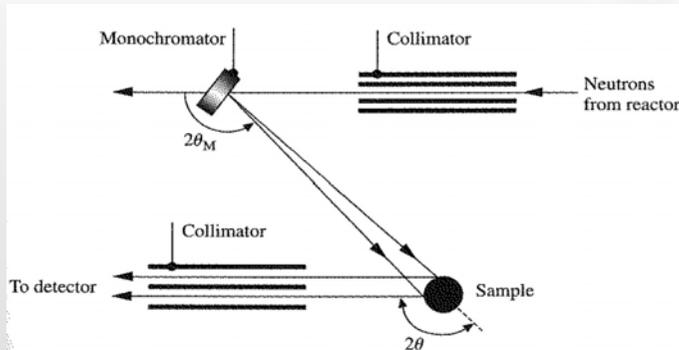


multi-wire counter

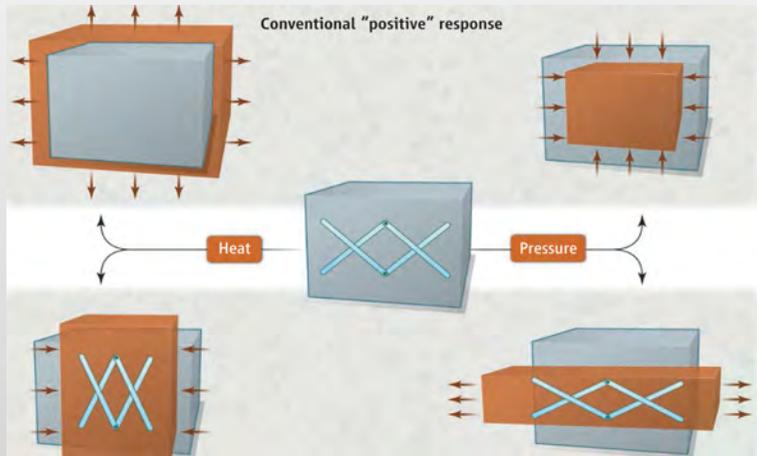
micro-pattern gas counter photolithography



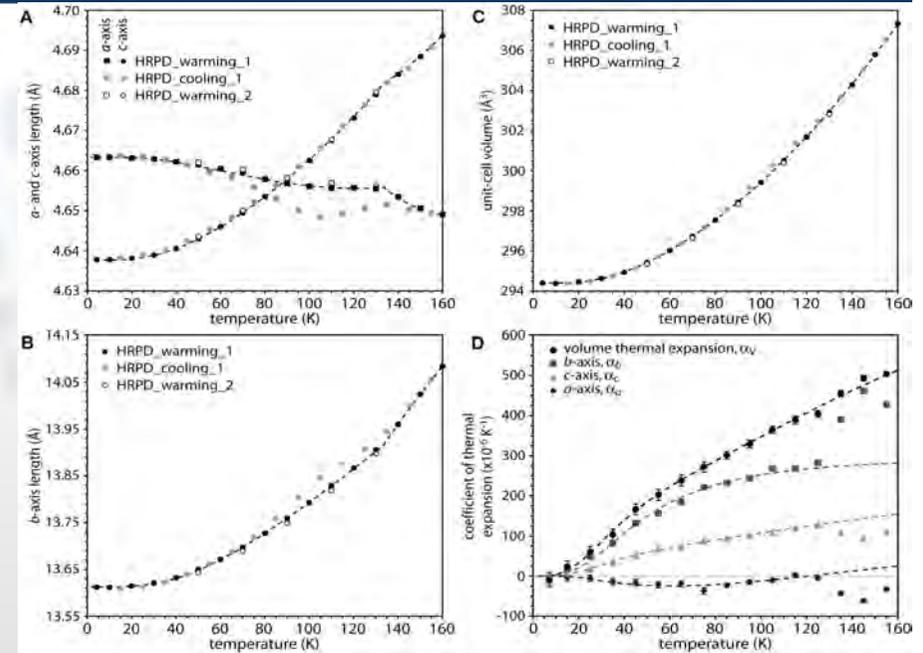
powder diffraction



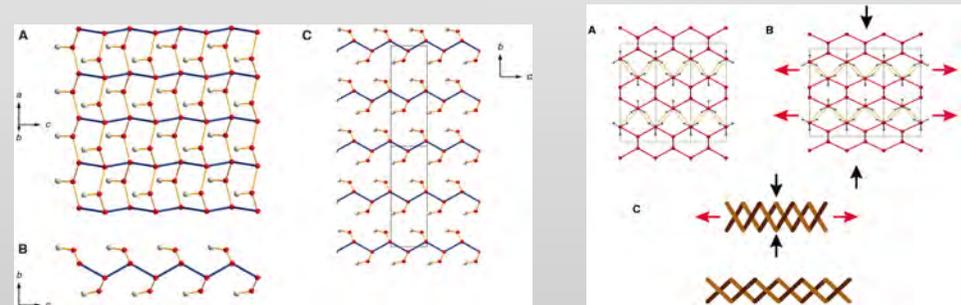
molecular crystal
 $\text{CH}_3\text{OH}\cdot\text{H}_2\text{O}$
 methanol monohydrate
 negative linear compressibility
 negative thermal expansion



applications to nano-switches
 but also planetary sciences ...



parametric studies of lattice dimensions



A D Fortes et al. Science 2011;331:742-746

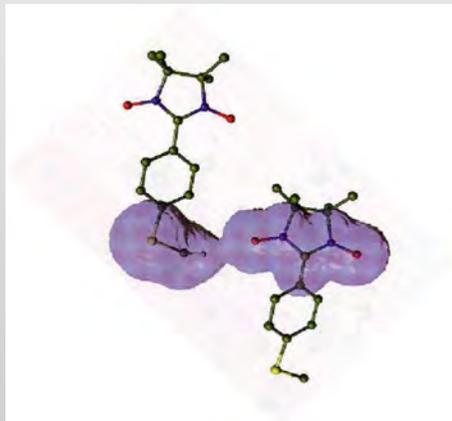
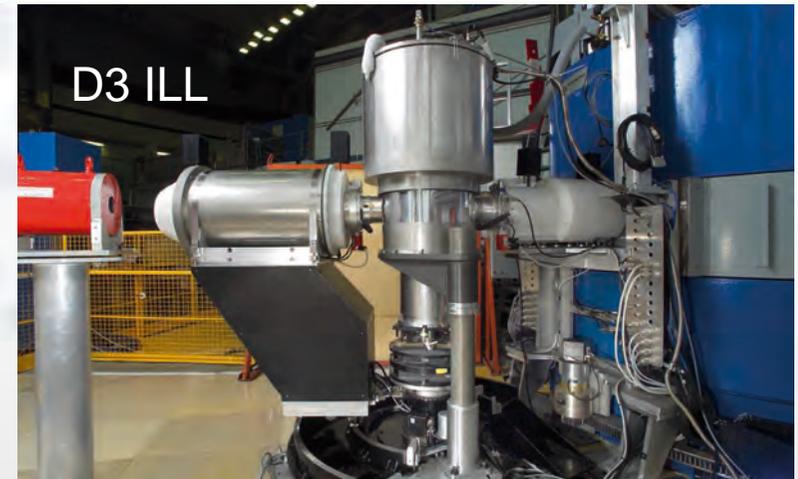
single crystal diffraction polarised neutrons

organic materials

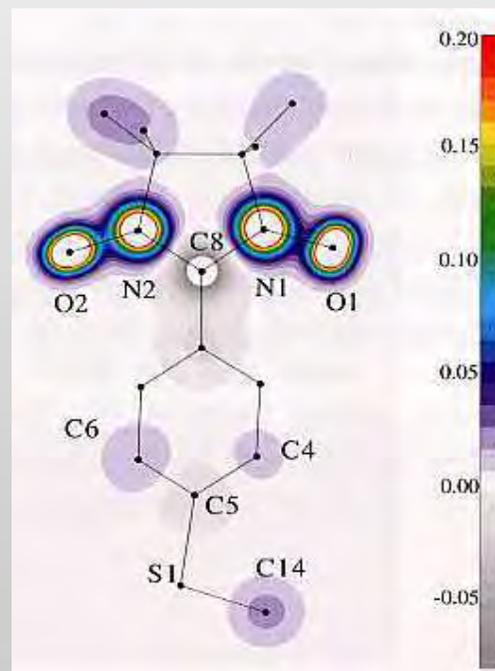
molecular magnets

'high' T_c

magnetic couplings driven by radicals packing

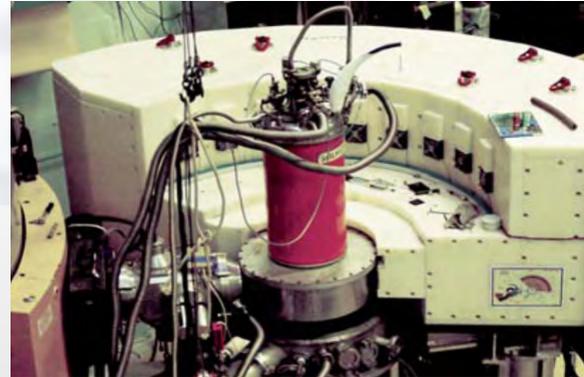
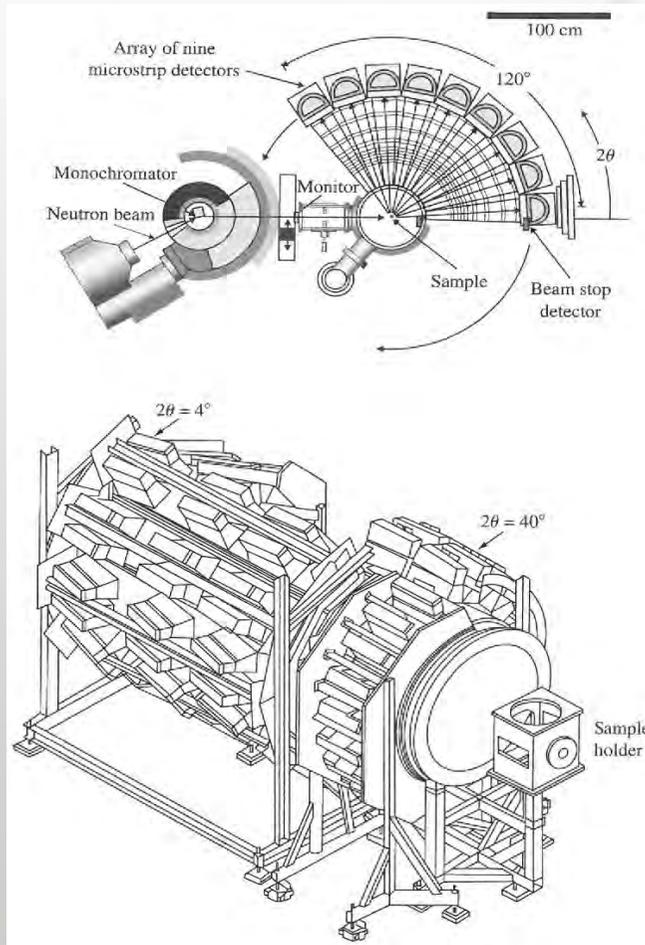


Nitrophenyl nitronyl nitroxide $T_c=0.6K$



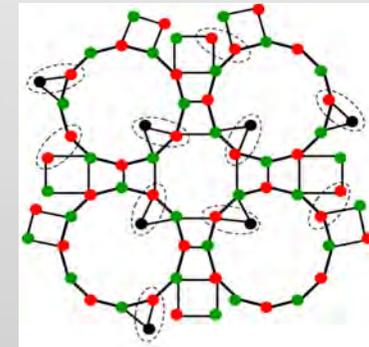
S. Pillet et al. (2001)

diffraction - disordered materials



D4 ILL

Sandals
ISIS

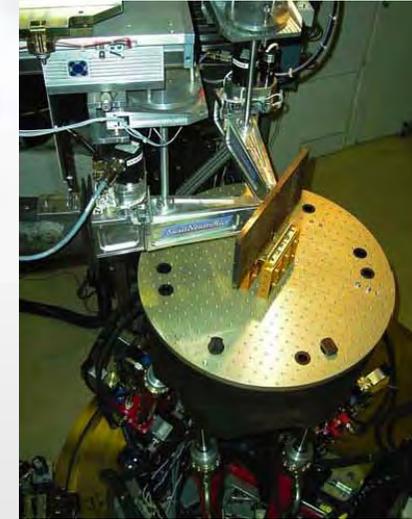


long range order observed in glasses
(chemical and topological)

diffraction and imaging engineering

residual stresses in large mechanical parts
welds,

SALSA instrument ILL



neutron imaging of large objects

fossil with leaves

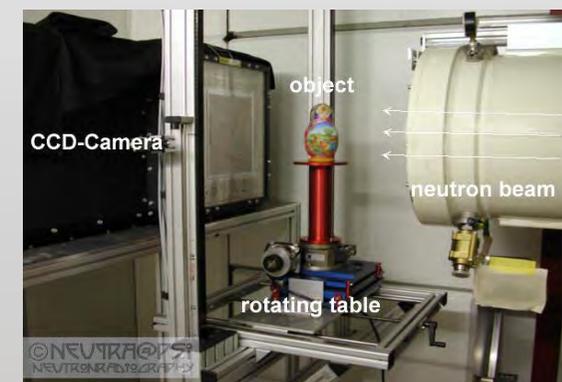


BMW engine



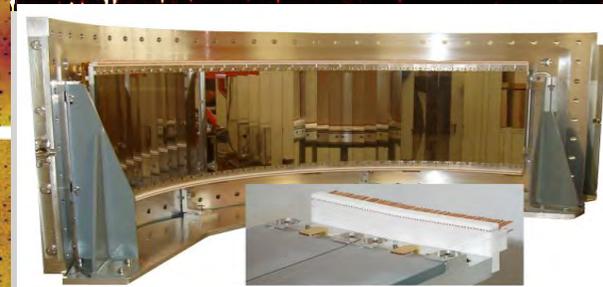
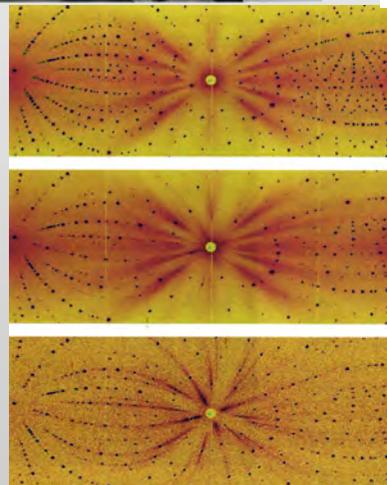
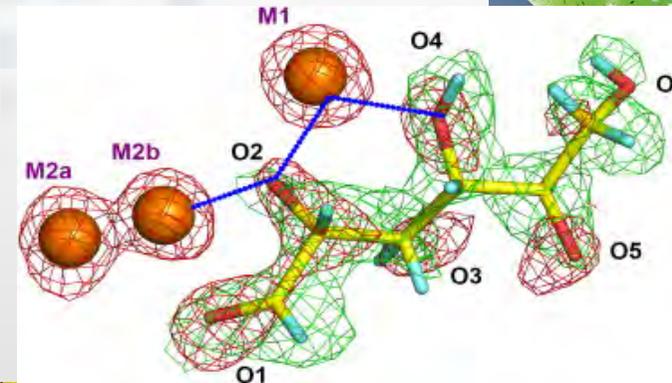
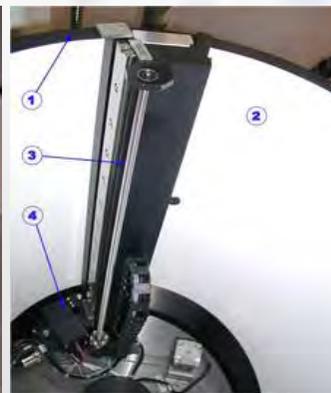
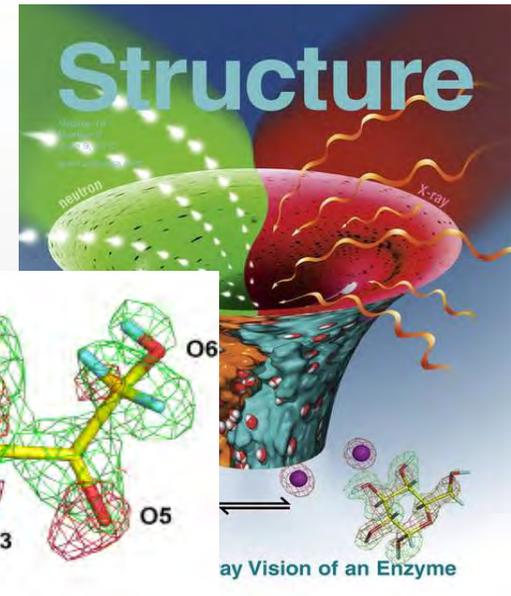
200 μ sec exposures

FRM-2 PSI ILL



diffraction and bio-molecules

deuteration complementarity SRX neutrons

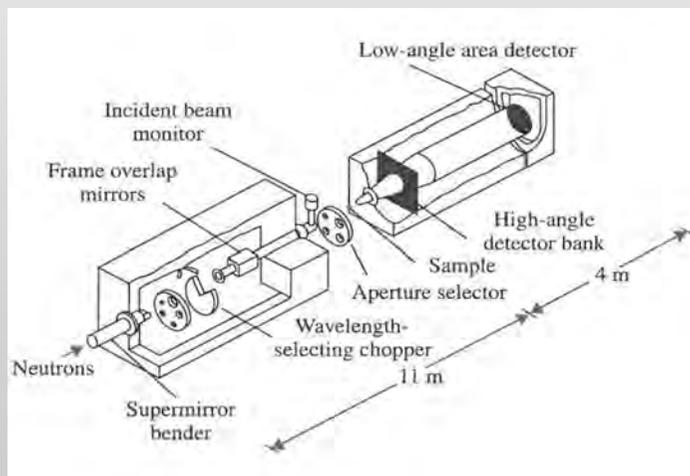
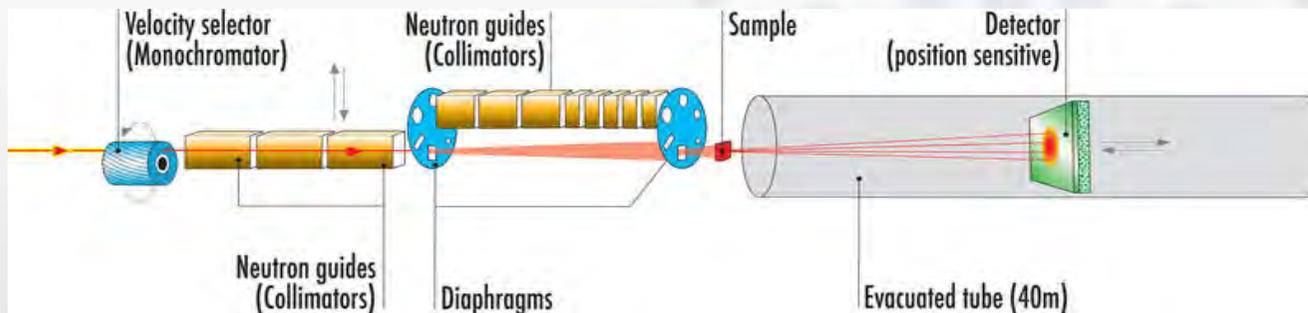


localisation and displacements of protons A.Y. Kovalevsky et al., Structure 18, 688 (2010)

instruments

small angle neutron scattering

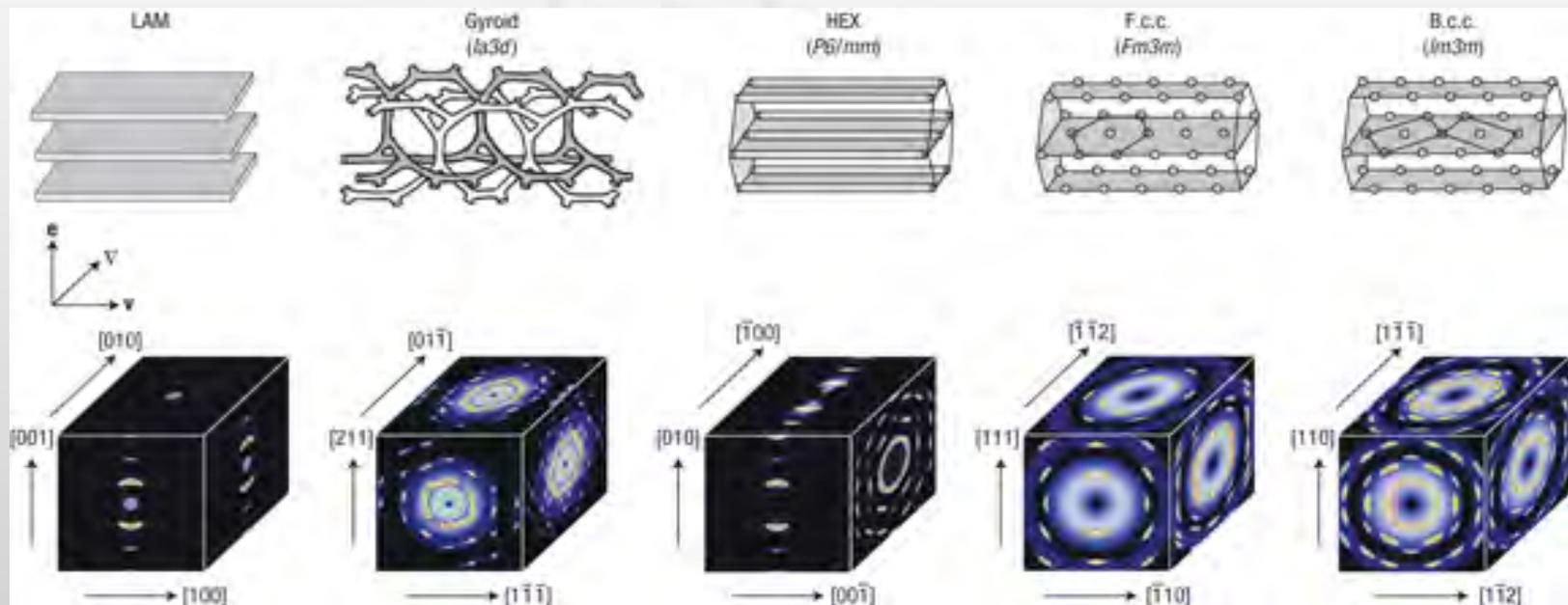
D11
ILL



LOQ
ISIS

'soft matter' systems

unified scheme of order and mutual orientation
under 'orientation' fields: flow, shear,

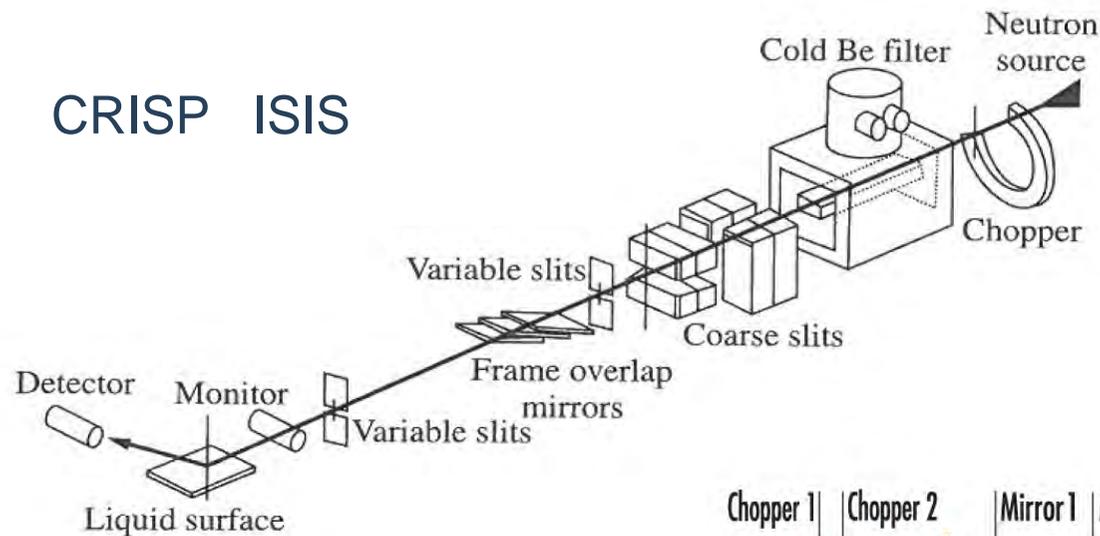


S. Förster et al. Nature Materials 6, 888 (2007)

instruments

neutron reflectometry

CRISP ISIS



Figaro ILL

